## **Impact of Higgsplosion on Relic Density of Dark Matter**

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S. Enomoto, N. Hiroshima, K. Murase, MY, arXiv:2310.XXXXX









### **Discovery of the Higgs boson**

Higgs properties

- -- origin of particle masses
- -- mass 125.11 GeV
- -- electrically neutral
- -- spin 0 (fundamental scalar particle)
- -- interaction with itself



Did the discovery complete the particle physics? Have we found all the laws of nature?

#### **Mysteries to be unraveled**



J. Ellis, M. Gaillard, D. Nanopoulos, arXiv:1504.07217

What is the origin of symmetry breaking? How many Higgs fields? https://phys.org/news/2011-12-dark.html rotational velocity tkm/sj 000 000 000 distance from center (light years) 00000 distance from center (light years) 00000 distance from center (light years) 00000 Dark Energy 68.3%

> What is dark matter (DM)? How was it generated?

Find the fundamental model describing the Higgs and DM in a unified picture!

So many literatures suggest that the Higgs is a bridge between the DM and our world

Important and necessary to carefully investigate the connection between the Higgs and DM



#### **High-multiplicity scalar production**

J. M. Cornwall, PLB243 (1990) H. Goldberg, PLB 246 (1990)

#### Higgs property

- -- origin of particle masses
- -- mass 125.25 GeV ----
- -- spin 0 (fundamental scalar particle)
- -- interaction with itself



Exponential growth of the "decay rate" of energetic particle with final state multiplicity

$$\Gamma_{\rm n} \sim \lambda^n n! \times f_n(E)$$

Contribution of large number of diagrams to the amplitude (no destructive interference)

### Higgsplosion

V. Khoze, M. Spannowski, NPB 926 (2018)

"Cross section" for 1 energetic Higgs to *n*-Higgs

$$\mathcal{R}_{n}(s) \equiv \frac{1}{2m_{\phi}^{2}} \frac{1}{n!} \int \frac{d^{3}p_{1}}{(2\pi)^{3}} \dots \frac{d^{3}p_{n}}{(2\pi)^{3}} \frac{1}{2E_{1}\dots 2E_{n}} (2\pi)^{4} \delta^{(4)} \left(q_{\phi} - p_{1}\dots - p_{n}\right) \left|\mathcal{M}(\phi^{*} \to n\phi)\right|$$
$$= \exp\left[n\left(\ln\frac{\lambda n}{4} + \frac{2}{\sqrt{3}}\frac{\Gamma(5/4)}{\Gamma(3/4)}\sqrt{\lambda n} - 1 + \frac{3}{2}\left(\ln\frac{\epsilon}{3\pi}\right) - \frac{25}{12}\epsilon\right)\right]$$

*n* dependence of *n*-Higgs production at proton-proton collider



### Aim



Indirect detection (current universe)

## Revisit the Higgs portal DM with taking into account high-multiplicity final state

-- precisely calculate the relic density to make use of a probe for DM-Higgs interaction

-- analyze the indirect signals of DM annihilation to reconstruct the nature of DM from cosmic rays



Talk plan

#### 1. Introduction

- 2. Setup and formulation
- 3. Numerical result
- 4. Summary and discussion

# **Setup and formulation**

#### Model

 $\phi$ : Higgs ( $\varphi$  after symmetry)

Note: Applicable to other models of a general scalar

Standard Model + dark matter  $\chi$ 

$$\mathcal{L} = \frac{1}{2} \left( \partial \phi \right)^2 - \frac{1}{4} \lambda \left( \phi^2 - v^2 \right)^2 + \bar{\chi} \left( i \partial - m_\chi \right) \chi - \left( y_\chi \phi \overline{\chi_R} \chi_L + \text{h.c.} \right)$$

$$\xrightarrow{\text{Symmetry}} \mathcal{L}_{\text{int}} = -\lambda v \varphi^3 - \frac{1}{4} \lambda \varphi^4 - \varphi \bar{\chi} \left( \tilde{y}_\chi P_L + \tilde{y}_\chi^* P_R \right) \chi \qquad \left( \begin{array}{c} \tilde{y}_\chi = y_\chi e^{-i \arg M_\chi} \\ M_\chi = m_\chi + y_\chi v \end{array} \right)$$

Symmetry breaking



#### **Transition amplitude**

Higgs "decay" into *n*-body Higgs  $\supset$  Higgsprosion effect



#### **Transition amplitude**

Higgs "decay" into *n*-body Higgs ⊃ Higgsprosion effect

$$\sum_{\text{spins}} \left| \mathcal{M}(\chi \bar{\chi} \to n\varphi) \right|^2 = \sum_{\text{spins}} \left| \mathcal{M}(\chi \bar{\chi} \to \varphi^*) \frac{1}{s - m_{\varphi}(s)^2 - im_{\varphi}(s)\Gamma_{\varphi}(s)} \mathcal{M}(\varphi^* \to n\varphi) \right|^2$$

DM annihilation to intermediate Higgs (straightforwardly calculated)

Dressed propagator ⊃ Higgspersion effect



**Dimensionless reaction rate** 

V. Khoze, M. Spannowski, NPB 926 (2018)

$$\mathcal{R}_{n}(s) \simeq \exp\left[n\left(L_{n} + \ln\frac{\lambda n}{4e} + \frac{3}{2}\ln\left(\frac{e}{3\pi}\frac{\sqrt{s} - nm_{\varphi}}{nm_{\varphi}}\right) - \frac{25}{12}\frac{\sqrt{s} - nm_{\varphi}}{nm_{\varphi}}\right)\right]$$

$$\stackrel{\frac{1}{s_{n} - m_{\varphi}^{2}}}{\stackrel{\frac{1}{s_{n} - m_{\varphi}^{2}}{\frac{1}{s_{n} - m_{\varphi}^{2}}}} n = \frac{\frac{1}{s_{n_{1}} - m_{\varphi}^{2}}\frac{i\mathcal{M}(1 \to n_{1})}{\sum_{i \neq 1}^{n} n_{i}}}{\sum_{i \neq 1}^{n} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\int_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}} n_{i}}{\sum_{i \neq 1}^{n} n_{i}} n_{i}} + \frac{\int_{i \neq 1}^{n} n_{i}} + \frac{\int_{i \neq$$

Ref: e.g., M. V. Libanov, V. A. Rubakov, D. T. Son, S. V. Troitsky, PRD50 (1994)

**Dimensionless reaction rate** 

V. Khoze, M. Spannowski, NPB 926 (2018)

$$\mathcal{R}_{n}(s) \simeq \exp\left[n\left(L_{n} + \ln\frac{\lambda n}{4e} + \frac{3}{2}\ln\left(\frac{e}{3\pi}\frac{\sqrt{s} - nm_{\varphi}}{nm_{\varphi}}\right) - \frac{25}{12}\frac{\sqrt{s} - nm_{\varphi}}{nm_{\varphi}}\right)\right]$$
  
From phase-space volume of *n*-Higgs final state  
Higher-order contribution V. Khoze, JHEP 06 (2017)  
$$L_{n} = \frac{2}{\sqrt{3}}\frac{\Gamma(5/4)}{\Gamma(3/4)}\sqrt{\lambda n} \simeq 0.854\sqrt{\lambda n}$$

**Important:** argument of the exponential = positive-valued

 $\rightarrow$   $R_n(s)$  grows with the multiplicity n

#### **Boltzmann equation** (evolution equation of DM density)

$$\begin{aligned} \frac{dn_{\chi}}{dt} + 3Hn_{\chi} &= -\int \frac{d^3k_{\chi}}{(2\pi)^3 2E_{\chi}} \frac{d^3k_{\bar{\chi}}}{(2\pi)^3 2E_{\bar{\chi}}} \Big[ f_{\chi} f_{\bar{\chi}} - f_{\varphi_1} \cdots f_{\varphi_n} \Big] \\ &\times \left| \tilde{y}_{\chi} \right|^2 \left( s - 4 |M_{\chi}|^2 \cos \theta_{\bar{y}_{\chi}} \right) \frac{1}{s^2 + m_{\varphi}^4 \mathcal{R}(s)^2} m_{\varphi}^2 \mathcal{R}_n(s) \\ & \left( \text{With Maxwell-Boltzmann distribution and energy conservation} \\ f_{\chi} f_{\bar{\chi}} - f_{\varphi_1} \cdots f_{\varphi_n} &= \frac{1}{(n_{\chi}^{eq})^2} \left[ (n_{\chi})^2 - (n_{\chi}^{eq})^2 \right] \\ &= - \left[ (n_{\chi})^2 - (n_{\chi}^{eq})^2 \right] \frac{1}{(n_{\chi}^{eq})^2} \frac{2 |\tilde{y}_{\chi}|^2 T^4}{(4\pi)^4} \\ & \times \int \frac{ds}{s} \frac{1}{T^3} \sqrt{s - 4M_{\chi}^2} \left( s - 4 |M_{\chi}|^2 \cos \theta_{\bar{y}_{\chi}} \right) K_1(\sqrt{s}/T] \underbrace{\frac{2m_{\varphi}^2 \mathcal{R}(s)}{s^2 + m_{\varphi}^4 \mathcal{R}(s)^2}}_{Window function W(s)} \end{aligned}$$

**Boltzmann equation** (evolution equation of DM density)

 $\frac{dn_{\chi}}{dt} + 3Hn_{\chi} = -\int \frac{d^3k_{\chi}}{(2\pi)^3 2E_{\chi}} \frac{d^3k_{\bar{\chi}}}{(2\pi)^3 2E_{\bar{\chi}}} \Big[ f_{\chi}f_{\bar{\chi}} - f_{\varphi_1} \cdots f_{\varphi_n} \Big]$ Final state multiplicity depends only on the self-coupling  $\lambda$ 1 = 0.05= 0.20.5= 0.10.8 If  $2M_{\chi} > \sqrt{s_{peak}}$ , the final state does not explode (averaging integral covers the region outside W(s) only) 0.6 W(s)For the case of SM ( $\lambda \simeq 0.13$ ): 0.4  $\sqrt{s_{peak}} \simeq 195 m_{\omega}$  with  $\Delta \sqrt{s} \simeq \pm 1 m_{\omega}$ 0.2 0  $= -\left[\left(n_{\chi}\right)^{2} - \left(n_{\chi}^{eq}\right)^{2}\right] \frac{1}{\left(n_{\chi}^{eq}\right)^{2}} \frac{2\left|\tilde{y}_{\chi}\right|^{2}T^{4}}{(4\pi)^{4}}$ 100 200 300 500 0 400  $\sqrt{s}/m_{\varphi}$  $\times \int \frac{ds}{s} \frac{1}{T^3} \sqrt{s - 4M_{\chi}^2} \left(s - 4|M_{\chi}|^2 \cos \theta_{\tilde{y}_{\chi}}\right) K_1(\sqrt{s}/T) \frac{2m_{\varphi}^2 s \mathcal{R}(s)}{s^2 + m_{\varphi}^4 \mathcal{R}(s)^2}$  $= -\langle \sigma v \rangle \left[ \left( n_{\chi} \right)^2 - \left( n_{\chi}^{eq} \right)^2 \right]$ Window function W(s)

## **Numerical result**

#### Interaction rate/Hubble rate vs $M_{\chi}/T$

(rough criterion)

#### $\square \text{ Maximized by } 2|M_{\chi}| \simeq 192m_{\varphi}$

smaller compared with the expectation from window function,  $2|M_{\chi}| \simeq 195 m_{\varphi}$ , due to thermal kinetics of DM

#### **D** Small interaction rate for $2|M_{\chi}| < 190m_{\varphi}$

mismatch between the window function and other part in thermal averaging due to  $K_1(\sqrt{s}/T) \ll 1$  wherein the window opens

#### **Relic density**



□ Parameter set ( $M_{\chi} = 4.8$  TeV,  $\tilde{y}_{\chi} = 1.53i$ ) successfully accounts for relic abundance

Much heavier than the Higgs portal DM in previous works,  $m_{\rm DM} \simeq 62$  GeV, where relic density is achieved by the Higgs pole

 Quantum statistics for the highmultiplicity state could change the results

Bose-Einstein distribution should be applied for the thermal averaging of DM annihilation, which may be enhanced by stimulated emission

# **Summary and discussion**

### **Summary and discussion**

#### □ Revisit to Higgs portal DM with taking into account Higgsplosion

- -- energetic Higgs boson decays into *n*-Higgs boson
- -- long-stay in equilibrium through strong interaction with Higgsplosion
- -- a favored parameter:  $M_{\chi} = 4.8$  TeV and  $|\tilde{y}_{\chi}| = 1.53$

(much heavy compared with Higgs portal DM in previous works)

-- simple and applicable to various models



#### -- quantum statistics effects for high-multiplicity

Bose-Einstein distribution and stimulated emission could change the shape of window function W(s)

-- test in indirect search of DM

important and necessary to reanalyze the signal with high-multiplicity state to reconstruct the nature of DM

