

Effects of the creation of particles in the field of waves of high intensity

Dmitrieva E., Satunin P.
INR RAS,
MSU

- **Particle creation by a dynamic classical field. Example: the production of matter particles by the inflaton field oscillation at the end of inflation (reheating stage) [1]**
- **Particle production enhancement effect - parametric resonance**
- **Creation of massive particles by an intense wave of a massless field with resonant amplification [2], small mass $m \ll \omega$**
- **The purpose of the work is to consider the case of arbitrary masses**

[1] Rubakov V.A. Gorbunov D.S. Introduction to the theory of the early universe. Cosmological perturbations. Inflationary theory. URSS, 2009.

[2] A.Arza. Phys.Rev.D 105 (2022) arXiv:2009.03870

$\mathcal{H} = g\phi\chi^2$ Two-scalar model

Equations of motion

$$(\square + m_\phi^2)\phi = -g\chi^2$$

Monochromatic plane wave

$$(\square + m_\chi^2)\chi = -2g\phi\chi$$

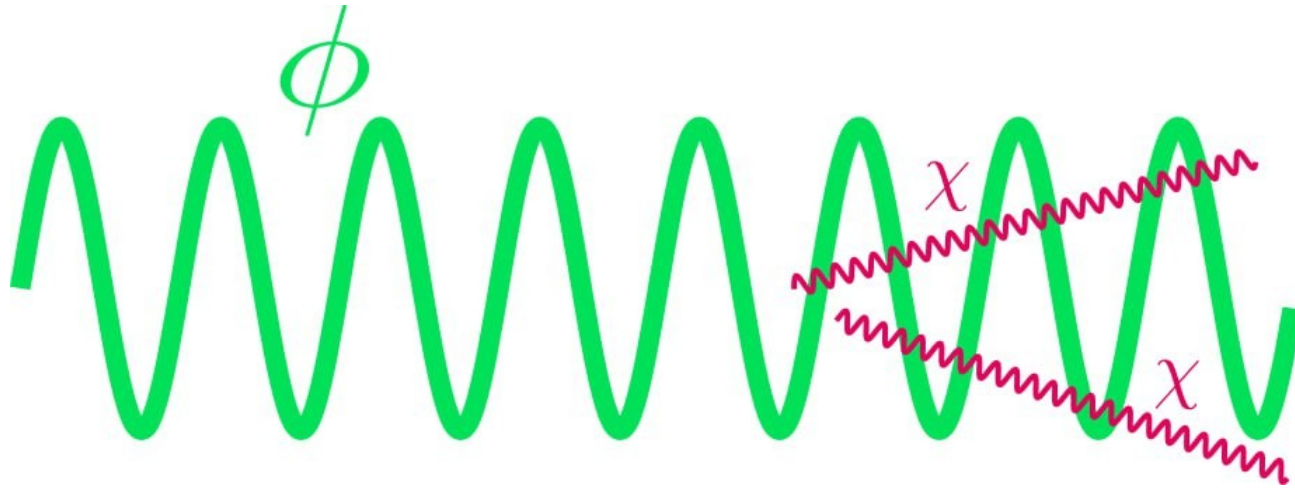
$$\phi(\vec{x}, t) = \frac{\sqrt{2\rho_\phi}}{\omega_p} \cos(\vec{p} \cdot \vec{x} - \omega_p t)$$

$$\Omega_{\vec{k}} = \sqrt{k^2 + m_\chi^2} \quad ; \quad \omega_{\vec{p}} = \sqrt{p^2 + m_\phi^2}$$

Quantum field

$$\chi = \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2\Omega_{\vec{k}}}} \left(\chi_{\vec{k}}(t) e^{i\vec{k} \cdot \vec{x}} + \chi_{\vec{k}}(t)^\dagger e^{-i\vec{k} \cdot \vec{x}} \right)$$

$$\mathcal{H} = g\phi\chi^2$$



Denote

$$A_{\vec{k}} = \chi_{\vec{k}} + \chi_{-\vec{k}}^\dagger$$

$$(\partial_t^2 + \Omega_{\vec{k}}^2)A_{\vec{k}} = -\omega_{\vec{p}}^2\alpha \left(\sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}}A_{\vec{k}-\vec{p}}e^{-i\omega_{\vec{p}}t} + \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}}A_{\vec{k}+\vec{p}}e^{i\omega_{\vec{p}}t} \right)$$

where $\alpha \equiv \frac{g\sqrt{2\rho_\phi}}{\omega_{\vec{p}}^3}$

Suppose $\chi_{\vec{k}} = a_{\vec{k}}(t)e^{-i\Omega_{\vec{k}}t}$

$$\begin{aligned}
& \frac{e^{-i\Omega_{\vec{k}}t}(\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}}\dot{a}_{\vec{k}}) + e^{i\Omega_{-\vec{k}}t}(\ddot{a}_{-\vec{k}}^\dagger + 2i\Omega_{-\vec{k}}\dot{a}_{-\vec{k}}^\dagger)}{=} \\
& = -\omega_{\vec{p}}^2\alpha \left(\sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}+\vec{p}}}} \left(a_{\vec{k}+\vec{p}} e^{-i(\Omega_{\vec{k}+\vec{p}}-\omega_{\vec{p}})t} + \underline{a_{-\vec{k}-\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}-\vec{p}}+\omega_{\vec{p}})t}} \right) + \right. \\
& \quad \left. + \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}} \left(a_{\vec{k}-\vec{p}} e^{-i(\Omega_{\vec{k}-\vec{p}}+\omega_{\vec{p}})t} + \underline{a_{-\vec{k}+\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}+\vec{p}}-\omega_{\vec{p}})t}} \right) \right)
\end{aligned}$$

The amplitude $a_{-\vec{k}+\vec{p}}^\dagger$ - leading, $a_{-\vec{k}-\vec{p}}^\dagger$ - subleading

The final equation

- **Approximation (A.Arza, PRD 2022):** \ddot{a} term neglected \rightarrow

$$m_\chi \ll \omega_p$$

$$\dot{a}_{\vec{k}} = -i\sigma_{\vec{k}} a_{\vec{p}-\vec{k}}^\dagger e^{i\epsilon_{\vec{k}} t}$$

$$\sigma_{\vec{k}} = g \sqrt{\frac{\rho_\phi/2}{\omega_{\vec{p}}^2 \Omega_{\vec{p}-\vec{k}} \Omega_{\vec{k}}}}$$

- **Without approximation**

$$e^{-i\Omega_{\vec{k}} t} (\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}} \dot{a}_{\vec{k}}) = \sigma_{p-k} a_{-\vec{k}+\vec{p}}^\dagger e^{i(\Omega_{-\vec{k}+\vec{p}} - \omega_{\vec{p}}) t}$$

$$\sigma_{p-k} = -\omega_p^2 \alpha \sqrt{\frac{\Omega_{\vec{k}}}{\Omega_{\vec{k}-\vec{p}}}}$$

Solution

- **In the approximation**

$$\epsilon_{p-k} = \Omega_k + \Omega_{p-k} - \omega_p = \epsilon_k$$

$$a_k(t) = e^{i\epsilon_k t/2} \left[a_k(0) \left(\cosh(s_k t) - i \frac{\epsilon_k}{2s_k} \sinh(s_k t) \right) - i \frac{\sigma_k}{s_k} a_{p-k}^\dagger(0) \sinh(s_k t) \right]$$

- **Without approximation**

$$s_{\vec{k}} = \sqrt{\sigma_{\vec{k}}^2 - \epsilon_{\vec{k}}^2/4}$$

$$a_k(t) = e^{i\epsilon_{p-k} t/2} \left[a_k(0) \left(\cosh(st) - i \frac{\frac{\epsilon_{p-k}^2}{4} - s^2 - \Omega_k \epsilon_{p-k}}{s(\epsilon_{p-k} - 2\Omega_k)} \sinh(st) \right) - i \frac{\sigma_{p-k}}{s(\epsilon_{p-k} - 2\Omega_k)} a_{p-k}^\dagger(0) \sinh(st) \right]$$

$$s_{p-k} = \sqrt{-\frac{\epsilon_{p-k}^2}{4} - 2\Omega_k^2 + \epsilon_{p-k}\Omega_k \pm \sqrt{\Omega_{p-k}^2 \epsilon_{p-k}^2 + 4\Omega_k^4 + \sigma_{p-k}^2 - 4\epsilon_{p-k}\Omega_k^3}}$$

Conditions of instability

$$s^2 > 0$$

Dimensionless quantities:

$$\vec{\kappa} = \vec{k}/\omega_p \quad \vec{v} = \vec{p}/\omega_p \quad \beta_k = \sqrt{\kappa^2 + \mu^2} \quad \mu = m_\chi/\omega_p \quad \eta_{p-k} = s_{p-k}/\omega_p$$

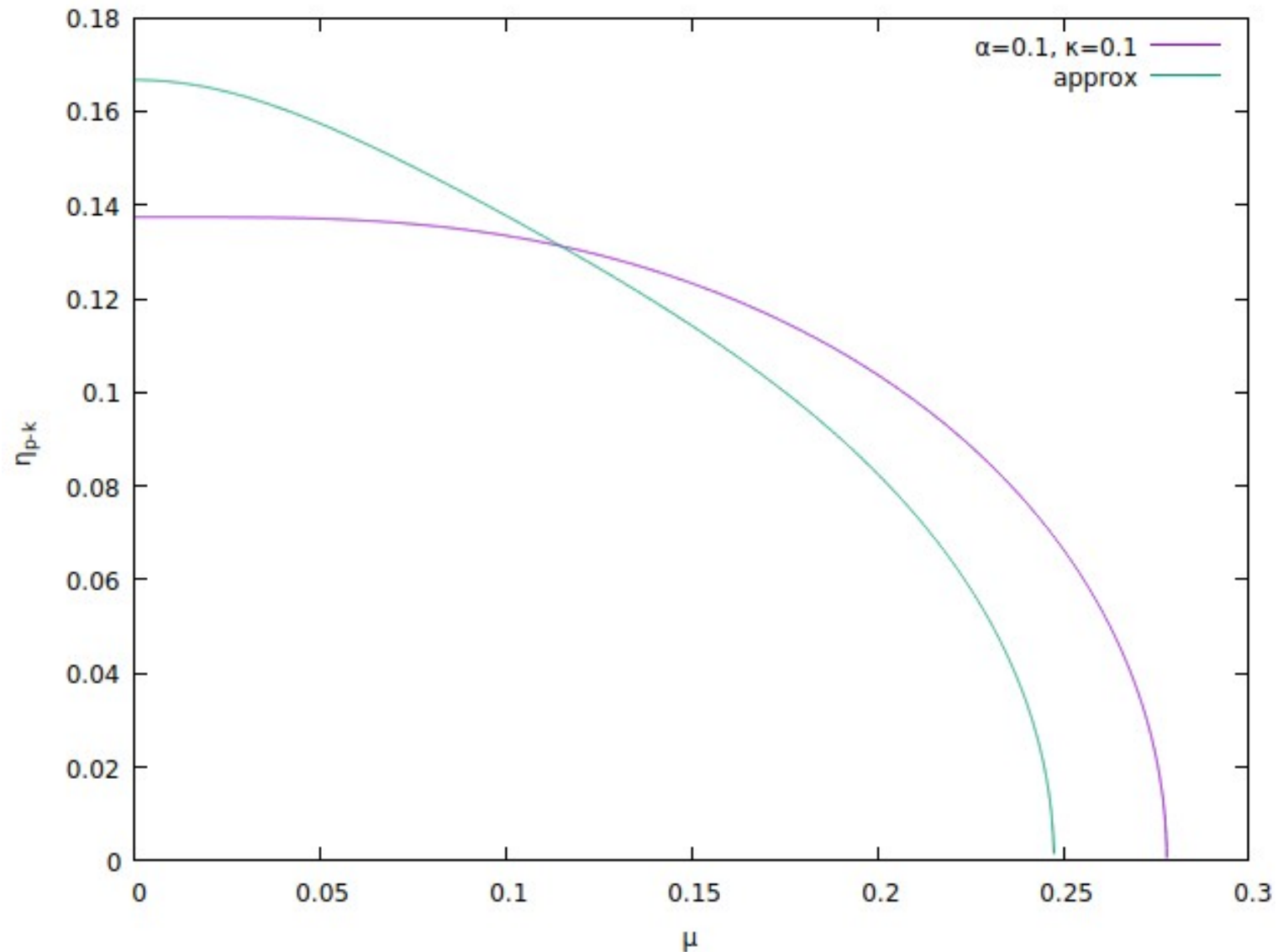
- In the approximation

$$\eta_{\vec{\kappa}} = \sqrt{\frac{\alpha^2}{4\beta_{\vec{v}-\vec{\kappa}}\beta_{\vec{\kappa}}} - \frac{1}{4}(\beta_{\vec{\kappa}} + \beta_{\vec{v}-\vec{\kappa}} - 1)^2}$$

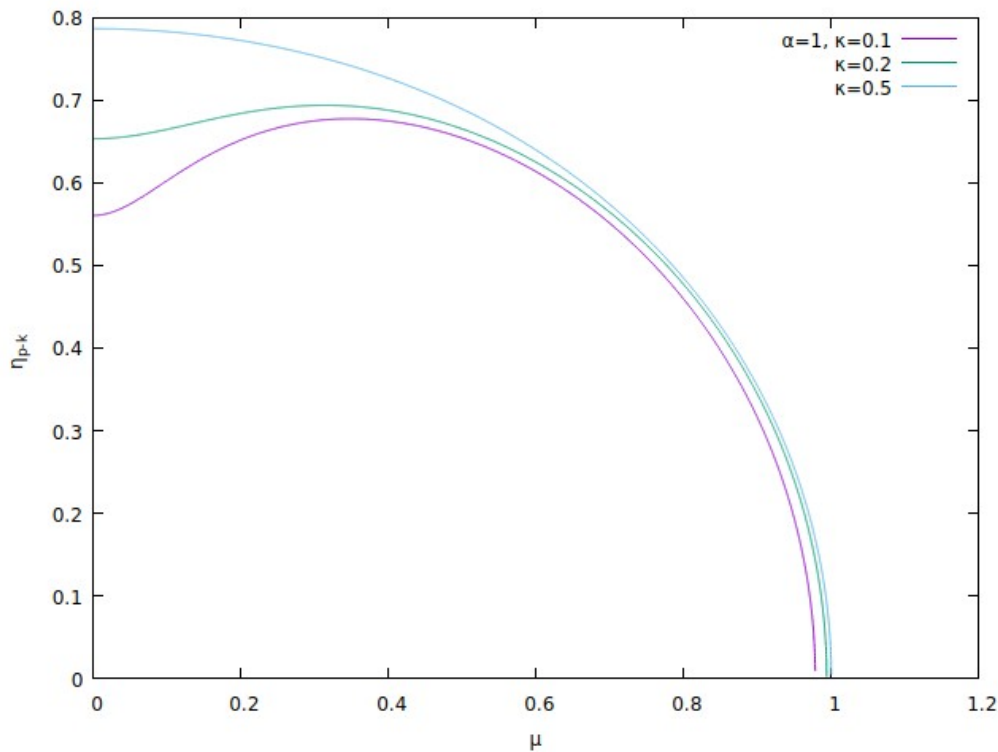
- Without approximation

$$\eta_{p-k} = \left(-\frac{1}{4}(\beta_k + \beta_{v-k} - 1)^2 - 2\beta_k^2 + (\beta_k + \beta_{v-k} - 1)\beta_k + \right. \\ \left. + \sqrt{\beta_k^2(\beta_k + \beta_{v-k} - 1)^2 + 4\beta_k^4 - 4(\beta_k + \beta_{v-k} - 1)\beta_k^3 + \frac{\beta_k\alpha^2}{\beta_{v-k}}} \right)^{\frac{1}{2}}$$

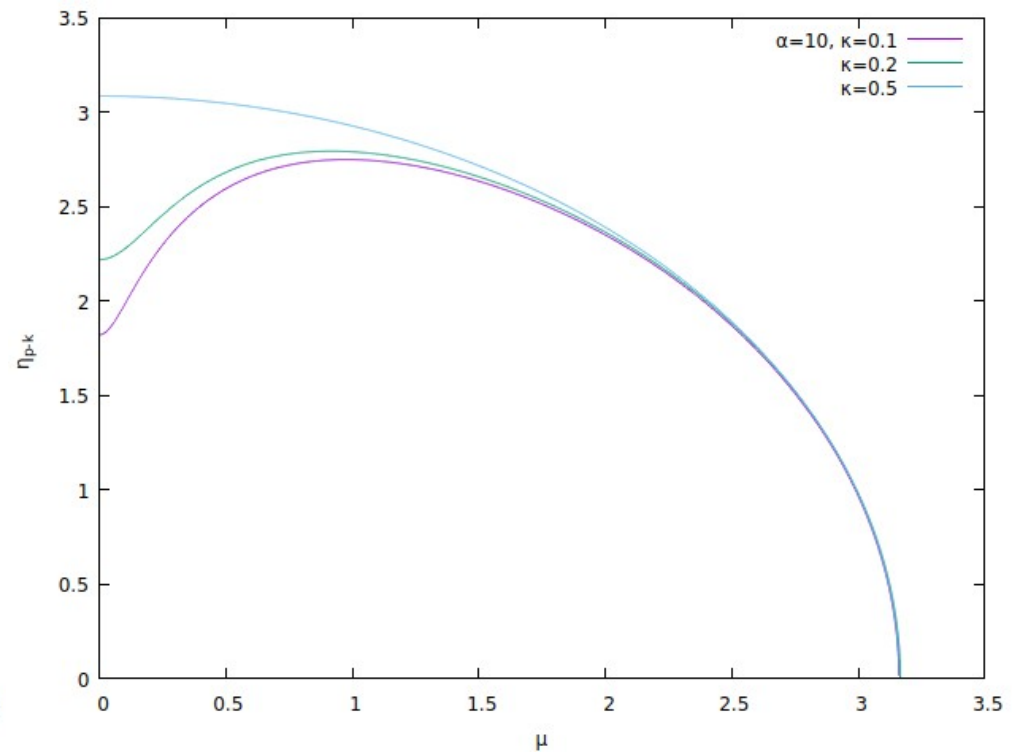
Dependence of η_{p-k} on μ at $\alpha=0.1$ and $\kappa=0.1$ in the approximation and without it



Dependences of η_{p-k} on μ for various α



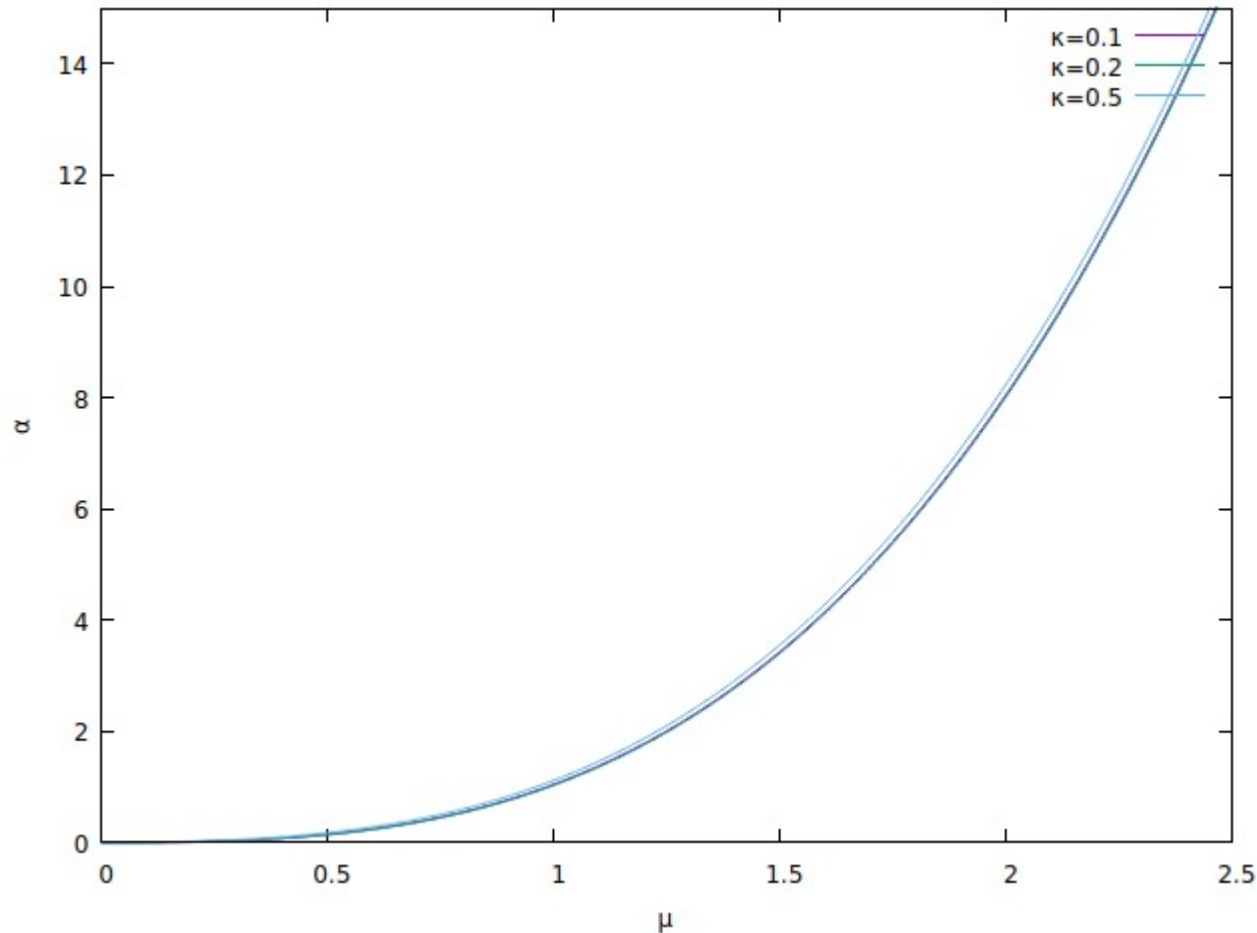
$\alpha=1$



$\alpha=10$

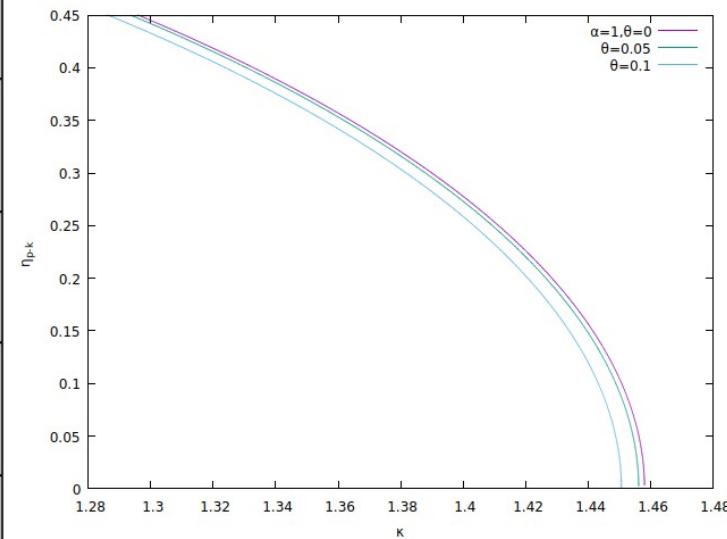
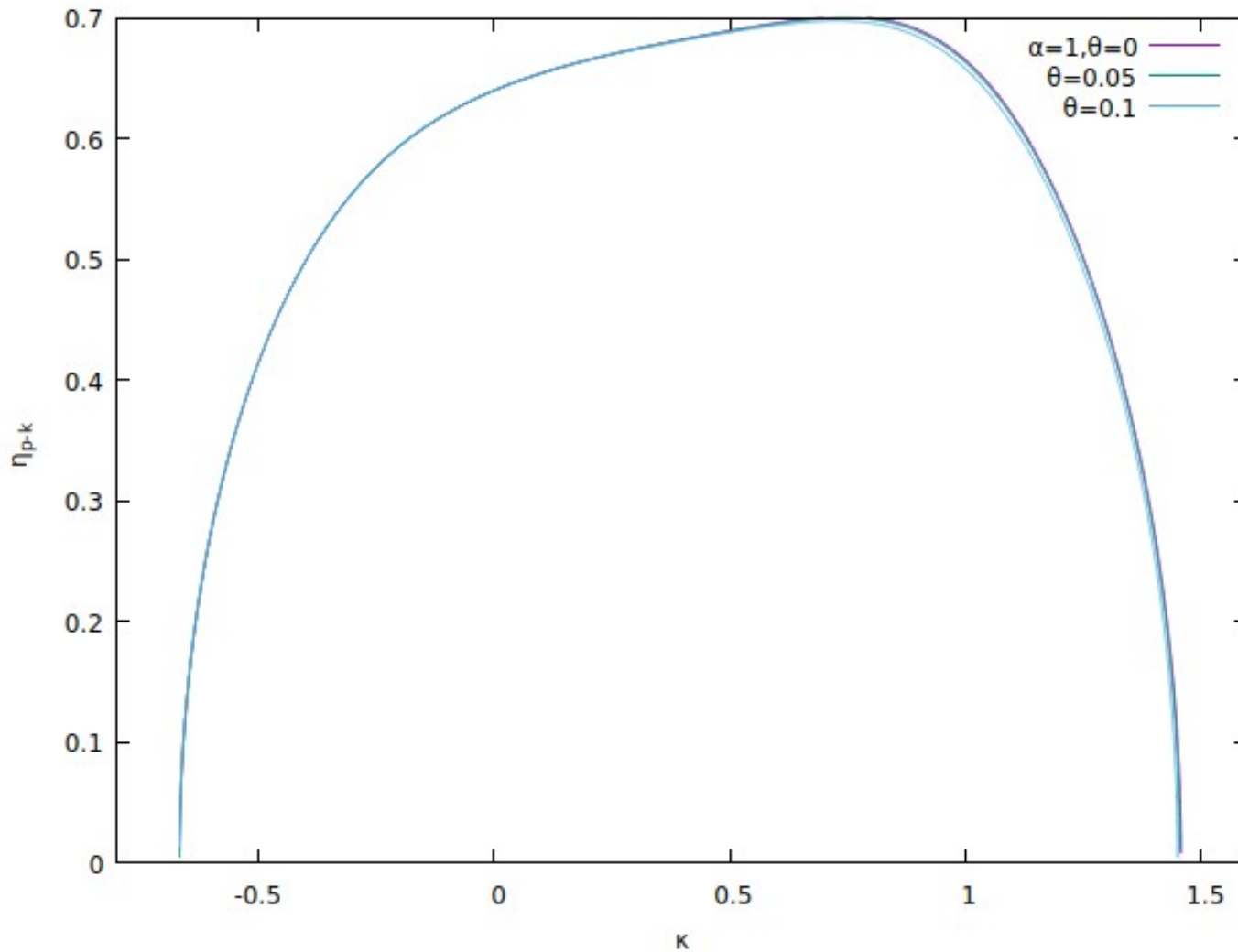
The boundary of stability

$$\eta_{p-k}^2 = 0$$



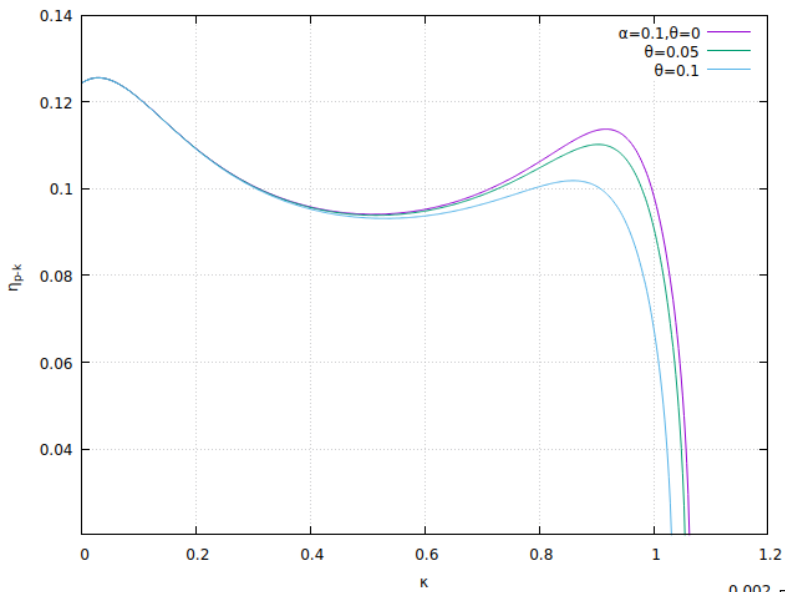
$$\alpha(\mu) = 4\sqrt{\frac{\beta_{v-k}}{\beta_k}(\beta_k + \beta_{v-k} - 1)^2(-3\beta_k + \beta_{v-k} - 1)^2}$$

Dependence of η_{p-k} on κ for $\alpha=1$



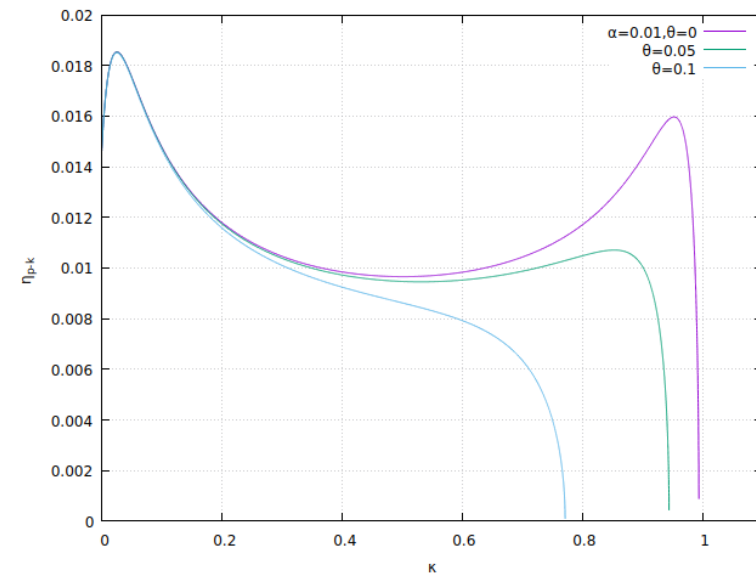
$$\mu = \sqrt{\alpha}/2$$

Dependence of η_{p-k} on κ for various α and fixed θ

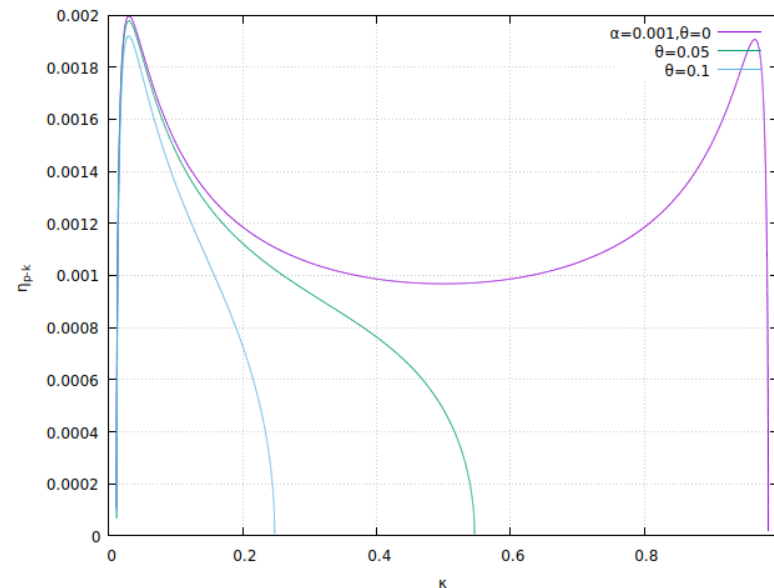


$\alpha=0.1$

$\alpha=0.001$



$\alpha=0.01$



Decay rate

The decay rate for the process $\phi_i \rightarrow \phi_j + \phi_k$ [3]

$$\Gamma = \int \frac{d^3 p_j}{(2\pi)^3} \int \frac{d^3 p_k}{(2\pi)^3} \frac{d|S_{fi}|^2}{dt}$$

Scattering matrix for this process

$$S_{fi} = \int_0^T dt \int d^3 x \langle f | H_I | i \rangle, \text{ where } H_I = g\phi\chi^2.$$

[3] N.Herring, B.Pardo, D.Boyanovsky arXiv:1808.02539

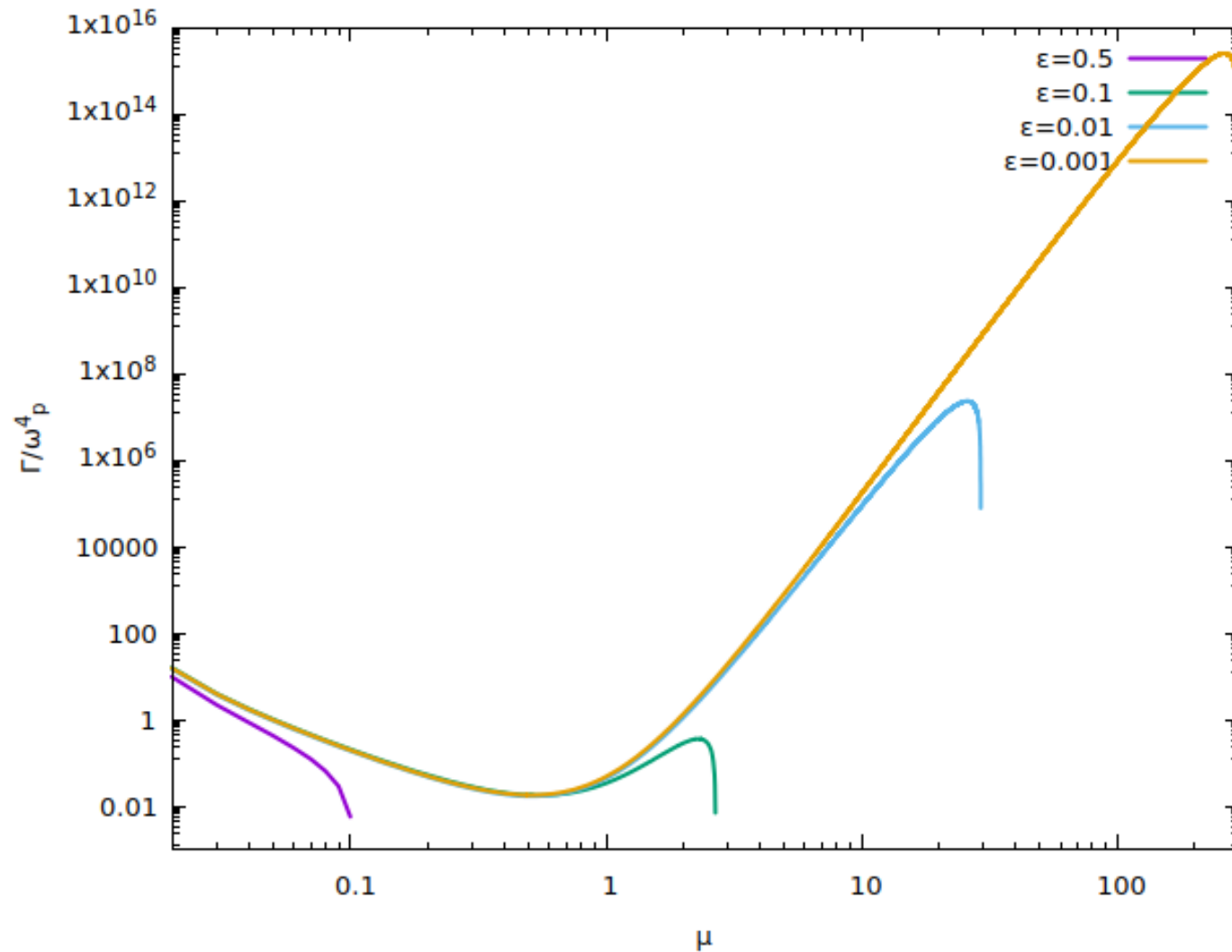
The final equation for the decay rate

$$\Gamma = \frac{\omega_{\vec{p}}^4 \alpha}{16(2\pi)^5} \int_0^\pi \int_{\kappa_-}^{\kappa_+} \frac{\sinh(\theta) \kappa^2 d\kappa d\theta}{\beta_{\vec{k}} \beta_{\vec{v}-\vec{k}}} \left(16C_1^2 \eta^2 (\sinh(2) - 4\eta_{\vec{v}-\vec{k}} \sinh(1)) + \frac{(1 - C_2^2 + C_1^2)^2}{4\eta_{\vec{v}-\vec{k}}} + \eta_{\vec{v}-\vec{k}} (1 - (C_2^2 - C_1^2)^2 (\cosh(1) + 1)) + 2\eta_{\vec{v}-\vec{k}}^3 \sinh(2) (1 + C_2^2 - C_1^2) \right)$$

where

$$C_1 = \frac{\frac{\epsilon_{\vec{v}-\vec{k}}^2}{4} - \eta_{\vec{v}-\vec{k}}^2 - \beta_{\vec{k}} \epsilon_{\vec{v}-\vec{k}}}{\eta_{\vec{v}-\vec{k}} (\epsilon_{\vec{v}-\vec{k}} - 2\beta_{\vec{k}})} \quad C_2 = \frac{\sigma_{\vec{v}-\vec{k}}}{\eta_{\vec{v}-\vec{k}} (\epsilon_{\vec{v}-\vec{k}} - 2\beta_{\vec{k}})}$$

Dependence of the decay rate on the μ for $\kappa_- = 0,1$ and $\kappa_+ = 0,9$



Standing wave

$$\phi(\xi, t) = 8 \frac{\sqrt{2\rho\phi}}{\omega} \cos(p_x \xi) \cos(p_y \xi) \cos(p_z \xi) \cos(\omega t)$$

The equation
$$e^{-i\Omega_{\vec{k}}t} (\ddot{a}_{\vec{k}} - 2i\Omega_{\vec{k}}\dot{a}_{\vec{k}}) = \sigma_{\vec{p}_j - \vec{k}} a_{\vec{p}_j - \vec{k}}^\dagger e^{i(\Omega_{\vec{p}_j - \vec{k}} - \omega)t}$$

$j = 1, \dots, 4$

Solution

$$a_{\vec{k}}(t) = e^{i\epsilon_{\vec{p}_j - \vec{k}}t/2} \left[a_{\vec{k}}(0) (\cosh(s_{\vec{p}_j - \vec{k}}t)) - i \frac{\frac{\epsilon_{\vec{p}_j - \vec{k}}^2}{4} - s_{\vec{p}_j - \vec{k}}^2 - \Omega_{\vec{k}}\epsilon_{\vec{p}_j - \vec{k}}}{s_{\vec{p}_j - \vec{k}}(\epsilon_{\vec{p}_j - \vec{k}} - 2\Omega_{\vec{k}})} \sinh(s_{\vec{p}_j - \vec{k}}t) - i \frac{\sigma_{\vec{p}_j - \vec{k}}}{s_{\vec{p}_j - \vec{k}}(\epsilon_{\vec{p}_j - \vec{k}} - 2\Omega_{\vec{k}})} a_{\vec{p}_j - \vec{k}}^\dagger(0) \sinh(s_{\vec{p}_j - \vec{k}}t) \right]$$

$$\vec{\xi} = (\xi_x, \xi_y, \xi_z), \vec{p}_1 = (p_x, p_y, p_z), \vec{p}_2 = (p_x, -p_y, p_z)$$

$$\vec{p}_3 = (-p_x, -p_y, p_z), \vec{p}_4 = (-p_x, -p_y, p_z)$$

Conclusions

- **The decay of $\varphi \rightarrow \chi\chi$ occurs at the amplitude of the field φ above the threshold, not only in the case of small masses ($m_\chi \ll \omega_p$) - the case of A.Arza, but also at an arbitrary mass**
- **In the case of large masses ($m_\chi \gg \omega_p$), the required threshold amplitude of the field φ is powerfully greater in comparison with the case of small amplitudes**

Thanks for your attention!